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# The Self-Energy of an Electron Bound in a Coulomb Field

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Abstract-Recent progress in calculation of the one-loop selfenergy of the electron bound in the Coulomb field is summarized. The relativistic multipole expansion is introduced. This expansion is based on a single assumption: except for the part of the time component of the electron four-momentum corresponding to the electron rest mass, the exchange of four-momentum between the virtual electron and photon can be treated perturbatively. For non Sstates and normalized difference  $n^3 \Delta E_n - \Delta E_1$  of the S-states this itself yields very accurate results after taking the method to the third order. For the ground state the perturbation treatment of the electron virtual states with very high three-momentum is to be avoided. For these states one can always rearrange the pertinent expression in such a way that free-particle approximation is allowed. Combination of the relativistic multipole expansion and free-particle approximation yields very accurate result after taking the method to the ninth order. These results are in very good agreement with the previous results obtained by the partial wave expansion and definitely exclude the possibility that the uncertainity in determination of the proton radius comes from the uncertainity in the calculation of the one-loop self-

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### I. Introduction

Recently found disagreement between the radius of proton deduced from a comparison of the theory [1], [2] and experiment [3], [4] of Lamb shifts in the ordinary and muonic hydrogen calls for a verification of correctness of both the theory and experiment [5]. Only when this is done, more interesting possibilities can be examined [5].

Among various contributions to the Lamb shift in ordinary hydrogen, the most important is the self-energy of the bound electron in the one-loop approximation. This effect causes the shift of 2S-1S transition of about 7.32 GHz. The difference between the proton radii found in [1] and [4] leads [6] to the difference of about 82.22 kHz for 2S-1S transition. Thus the self-energy effect should be known with accuracy significantly better than 1 part in  $10^5$ .

For years the self-energy for the atoms with low nuclear charges has been calculated via the series in powers of  $Z\alpha$ 

$$\Delta E = \frac{m\alpha(Z\alpha)^4}{\pi n^3 s^3} F(Z\alpha, n, l_j), \tag{1}$$

where [7], [8], [9], [10]

$$F(Z\alpha, n, l_j) = A_{41} \ln s(Z\alpha)^{-2} + A_{40} + A_{50}(Z\alpha) + (2)$$
$$+ (Z\alpha)^2 \left[ A_{62} \ln^2 s(Z\alpha)^{-2} + A_{61} \ln s(Z\alpha)^{-2} + A_{60} \right] +$$
$$+ (Z\alpha)^3 \left[ \ln s(Z\alpha)^{-2} A_{71} + A_{70} \right] + \dots,$$

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where s=1 in the non-recoil limit. The dominant part of the recoil effect is taken into account by replacing s=1 in Eqs. (1) and (2) by  $s=1+m_e/m_n$ . Here  $m_e/m_n$  is a ratio of the electron and nuclear masses [13]. The coefficients A,  $A=A(n,l_j)$ , are summarized e.g. in [1], [11]. n, l and j denote the principal, orbital and total momentum quantum numbers of the state under consideration.

Recently, another approach based on the partial wave expansion [12] was developed. For atoms with low nuclear charges this expansion converges very slowly. Nontheless, the most accurate results obtained so far for the hydrogen atom were obtained in [11] by means of this expansion. In that paper several millions of partial waves were considered. For each partial wave there is a three-dimensional integration to be performed numerically, for details see [11], [12].

It is convenient to write the self-energy effect on general S-state as

$$\Delta E_n = \frac{n^3 \Delta E_n - \Delta E_1}{n^3} + \frac{\Delta E_1}{n^3},\tag{3}$$

where the first and the second terms on the right member will be referred to as the state-dependent and the state-independent parts, respectively. In comparison with calculation of the stateindependent part of the S-states, calculation of the statedependent part of the S-states and non S-states is substantially simpler: the coefficients  $A_{41}$ ,  $A_{50}$  and  $A_{62}$  in Eq. (2) vanish and the estaimate obtained by means of series (2) truncated after  $A_{60}$  term is quite good, see below. However, the dominant part of the efect is state independent part of the S-states. The relative difference between the series (2) truncated after  $\alpha(Z\alpha)^6$  term and the numerical result in [11] for the ground states of the hydrogen is 3 parts in 106. In view of the complexity of the calculation of the  $A_{60}$  coefficient [8], there is no hope of achieving significantly better accuracy within purely perturbation approach. Clearly, such an approach is not sufficient any more.

# II. THE METHOD

The renormalized expression for the self-energy in the non-recoil limit reads (see [14] for notation used)

$$\Delta E = \langle O - \Delta m \rangle = \langle \psi | \gamma_0 (O - \Delta m) | \psi \rangle, \tag{4}$$

where  $\Delta m$  stands for the electromagnetic mass of the electron and the regularized mass operator O reads

$$O = \frac{\alpha}{\pi} \int_0^{\Lambda^2} d\lambda \int \frac{d^4 k_F}{(k^2 - \lambda)^2} \gamma_\mu \frac{1}{\gamma \cdot (\Pi - k) - m} \gamma_\mu.$$
 (5)

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The wave function  $\psi$  is a solution of the stationary Dirac equation with the energy E

$$(\gamma.\Pi - m)\psi = 0. \tag{6}$$

The components of physical momentum  $\Pi$  of the particle in the case of the Coulomb field read

$$\Pi = \left(E + \frac{Z\alpha}{R}, \vec{P}\right) \tag{7}$$

where  $\vec{P}$  is the canonical momentum.

Following [7] we multiply  $1/(\gamma.(\Pi-k)-m)$  in Eq. (5) by  $(\gamma \cdot (\Pi - k) + m)/(\gamma \cdot (\Pi - k) + m)$  from the right. Using Dirac equation and properties of  $\gamma$  matricies [14] we get

$$\langle O \rangle = -\frac{\alpha}{2\pi} \left\langle \gamma_{\mu} \left( G_4 \Pi_{\mu} - \frac{m}{2} G. \gamma \gamma_{\mu} \right) \right\rangle,$$
 (8)

where

$$G_{4,\nu} = (-4) \int_0^{\Lambda^2} d\lambda \int \frac{d^4k_F}{(k^2 - \lambda)^2} \frac{(1, k_\nu/m)}{k^2 - 2k \cdot \Pi + H}.$$
 (9)

Here the second-order Hamilton operator H [7] reads

$$H = (\gamma . \Pi + m)(\gamma . \Pi - m) = \Pi . \Pi - m^2 + \frac{1}{4} [\Pi_{\mu}, \Pi_{\nu}] [\gamma_{\mu}, \gamma_{\nu}].$$
(10)

Recently an expansion of the electron propagator has been suggested [15], [16], [17], [18]. This expansion is referred to as relativistic generalization of multipole expansion (RME). This expansion is based on a single assumption, namely that the four-momentum  $\Pi$  of the bound electron in the virtual states is dominated by the four-momentum  $\varepsilon$  of the electron at rest:

$$\frac{1}{k^2 - 2k \cdot \varepsilon + H - 2k \cdot (\Pi - \varepsilon)} = \frac{1}{z - \tilde{H}_0 - \lambda \tilde{H}_1}.$$
 (11)

Here the propagator is written in a generic form  $(z - \tilde{H})^{-1}$ , where  $H = H_0 + \lambda H_1$  is a generic Hamilton operator. Furthermore,  $\lambda$  is a formal perturbation parameter that is eventually set to one,

$$\varepsilon = (m, 0, 0, 0) \tag{12}$$

and

$$z = k^2 - 2k.\varepsilon, \, \tilde{H}_0 = -H, \, \lambda \tilde{H}_1 = 2k.(\Pi - \varepsilon). \tag{13}$$

We form the series

$$\Delta E = m \frac{\alpha}{\pi} (Z\alpha)^4 \sum_{v=1}^{\infty} F_v, \tag{14}$$

$$-2(Z\alpha)^{4}F_{v} = \sum_{t=0}^{v} \left\langle \gamma_{0}G_{4}^{2(v-t),t} + \gamma_{\mu}G_{4}^{2(v-t-1),t} \frac{(\Pi-\varepsilon)_{\mu}}{m} + \right.$$

$$\begin{array}{c} t=0 \\ +\gamma_0 G_0^{2(v-t),t} - \frac{1}{2} \gamma_\mu [G_0^{2(v-t-3),t}, \gamma_0 \gamma_\mu] + \frac{1}{2} \gamma_\mu G_i^{2(v-t)-1,t} \gamma_i \gamma_\mu \\ \end{array} \\ \begin{array}{c} S_v = F_v - (Z\alpha) A_{50}^{(v)}. \end{array} \end{aligned}$$

The superscripts v and t in  $G^{v,t}_{4,\nu}$  stand for the number of expanded space and time components of  $(\Pi-\varepsilon)$ , respectively. The method for generating the individual terms of the expansion is described in [15], [18] for the ground state and in [16] for non S-states and state-dependent part of the S-states.

RME has a number of advantages. Once the renormalization of the electron mass is made all the integrals over either photon or electron variables are finite both at the infrared and ultraviolet. Thus, no separation of any of the integration is necessary. The terms of RME are generated very easily. In fact, by means of computer languages for symbolic calculation like Maple or Mathematica they can be generated automatically. The only integrals to be performed numerically are onedimensional integrals over the electron wave numbers of the continuous part of the spectrum. These integrals converge very fast. For non S-states and state-dependent part of the S-states it is sufficient to consider the sum in Eq. (14) only up to v=3. Further, for the atoms with low nuclear charges the pertinent expressions can be further simplified and part of the calculation can be carried out analytically, for details see [16].

Question then arises of the relation of the series (14) to (2). As argued in [15], [17], [16] the coefficients  $A_{41}$ ,  $A_{40}$ and  $A_{62}$  are contained in  $F_1 + F_2$ . For non S-states and state-dependent part of the S-states the coefficient  $A_{61}$  is contained in  $F_1 + F_2$  and the coefficient  $A_{60}$  is contained in  $F_1 + F_2 + F_3$ . For state-independent part of the S-states the coefficient  $A_{61}$  and sufficiently great part of the coefficient  $A_{60}$  is contained in  $F_1 + F_2 + F_3$ , see Tables 2 and 3 of [17]. The coefficient  $A_{50}$  is contained in complete sum (14). This coefficient is entirely determined by the virtual states with very high wave numbers. For these states one can always rearrange the pertinent expression in such a way that freeparticle approximation is appropriate, see [18]. The coefficient  $A_{50}$  can then be written as

$$A_{50} = \sum_{v=1}^{\infty} A_{50}^{(v)},\tag{16}$$

where  $A_{50}^{(v)}$  is the part of  $A_{50}$  coefficient contained in  $F_v$  [17],

$$A_{50}^{(v)} = -2^3 \frac{\Gamma\left(\frac{1}{2}\right)\Gamma\left(v - \frac{5}{2}\right)\left(16v^4 - 32v^3 + 296v^2 + 8v - 267\right)}{\pi\Gamma(v)(2v+5)(2v+3)(2v+1)^2(2v-3)}$$
(17)

For large v this behaves as

$$A_{50}^{(v)} \simeq -\frac{4}{\pi^{1/2}} v^{-7/2}.$$
 (18)

The complete coefficient  $A_{50}$  is [7], [8]

$$A_{50} = 4\pi \left(\frac{139}{128} - \frac{\ln 2}{2}\right). \tag{19}$$

For the self-energy function  $F(Z\alpha, 1, 0)$ , Eq. (1), we write

$$F(Z\alpha, 1, 0) = (Z\alpha)A_{50} + \sum_{v=1}^{\infty} S_v,$$
 (20)

$$S_v = F_v - (Z\alpha) A_{50}^{(v)}. (21)$$

particle result is contained in it. When considering the selfenergy effect of the light hydrogen-like ions this is the greatest advantage of the RME over the method used in [11], [12] based on partial wave expansion.

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## III. RESULTS

The results obtained for the low lying states of the hydrogen atom are displayed in Tables I and II taken from [16] and [18], respectively.

### TABLE I

The result for the scaled energy function  $F(\alpha,n,l_j)$  for low LYING S- AND P-STATES OF HYDROGEN-LIKE ATOMS OBTAINED BY VARIOUS METHODS. FOR THE S-STATES THE RESULTS FOR THE DIFFERENCE  $F(\alpha,n,0)-F(\alpha,1,0)$  are given. The results PRESENTED AS 'RME' WERE OBTAINED FROM THE SERIES (14), TRUNCATED AFTER THE THIRD TERM, FOR DETAILS SEE [16]. THE RESULTS OBTAINED BY MEANS OF PARTIAL WAVE EXPANSION ARE TAKEN FROM [11]. THE RESULTS OBTAINED BY MEANS OF Eq. (2) ARE TAKEN FROM [10].

State	RME	Partial waves	Eq. (2)
2s-1s	0.23002947	0.23003154	0.23003040
$2p_{1/2}$	-0.1263965908	-0.12639637	-0.12639773
$2p_{3/2}$	0.1234982498	0.12349856	0.12349766
3s-1s	0.28881828	0.28882057	0.28881946
4s-1s	0.31259262	0.31259475	0.31259391

### TABLE II

THE CONVERGENCE OF RME FOR THE GROUND STATE OF THE HYDROGEN. 'LEAD' STANDS FOR THE

HYDROGEN. 'LEAD' STANDS FOR THE 
$$\begin{aligned} & \text{SUM } S_1 + S_2 + A_{50}. \text{ Instead of Eq. (15) we considered } -2(Z\alpha)^4 F_v \simeq \\ & \left\langle \gamma_0 \sum_{t=0}^T \left( G_4^{2(v-t),t} + G_0^{2(v-t),t} \right) + \sum_{t=0}^{T-1} \left( \gamma_\mu G_4^{2(v-t-1),t} \frac{(\Pi-\varepsilon)_\mu}{m} + \frac{1}{2} \gamma_\mu \sum_{t=0}^{T-1} G_i^{2(v-t)-1,t} \gamma_i \gamma_\mu \right) \right\rangle, \text{ where } T = 3. \text{ The contribution } \\ & \text{ of the terms} < \gamma_0 G_{4,0}^{2v,t} > \text{ for } t > 3 \text{ the terms} \\ & < \gamma_\mu G_4^{2(v-t-1),t} (\Pi-\varepsilon)_\mu > \text{ and } < \gamma_\mu G_i^{2(v-t)-1,t} \gamma_i \gamma_\mu > \text{ for } t > 2 \\ & \text{ and } \left( -\frac{1}{2} \right) \left\langle \gamma_\mu [G_0^{2v,t}, \gamma_0 \gamma_\mu] \right\rangle \text{ are very small and are estimated } \\ & \text{ under heading 'small'. We have found that for large $v$ the ratio $S_v/S_{v-1}$ approaches the ratio $A_{50}^{(v)}/A_{50}^{(v-1)}$. This enabled us to estimate the remainder of the series, 'rem'. For details see [18]. 'Other' is the result of the partial wave expansion$$

SEE [18]. 'OTHER' IS THE RESULT OF THE PARTIAL WAVE EXPANSION TAKEN FROM [11].

term	$F(\alpha, 1, 0)$
lead	10.315870916
$S_3$	$0.891183 \ 10^{-3}$
$S_4$	$0.23509 \ 10^{-4}$
$S_5$	$0.4484 \ 10^{-5}$
$S_6$	$0.1554 \ 10^{-5}$
$S_7$	$0.719 \ 10^{-6}$
$S_8$	$0.388 \ 10^{-6}$
$S_9$	$0.237 \ 10^{-6}$
sum	10.3167929917
small	$0.36 \ 10^{-7}$
rem	$0.648 \ 10^{-6}$
total	10.316793675(50)
other	10.316793650(1)

These results agree with the results obtained by partial wave expansion to high degree of accuracy. The difference between the result for the state dependent and state independent part of the S-states obtained by these two completely independent methods amounts to the difference of 210.5 and 18 Hz for 2S-1S transitions in hydrogen, respectively. This definitely excludes the possibility that uncertainity in determination of the proton radius comes from the uncertainity in determination the one-loop self-energy correction to the energy levels.

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- The following notation is used:  $d^4k_F$  stands for  $i(2\pi)^{-2}d^4k$  and k stands for the four-momentum of the virtual photon. We use the summation convention  $\vec{A} \cdot \vec{B} = A_{\nu}B_{\nu} = A_{0}B_{0} - \vec{A} \cdot \vec{B}$ . In particular, summation convention  $A.B = A_{\nu}B_{\nu} - A_{0}B_{0}$  and  $k^{2} = k^{2}_{0} - \omega^{2}$ ,  $\omega^{2} = \vec{k}.\vec{k}$ . The Dirac  $\gamma$  matrices are defined by the anticommutation relation  $\{\gamma_{\mu},\gamma_{\nu}\}=2g_{\mu\nu}$ , where  $g_{0\mu}=\delta_{0\mu}$  and  $g_{ij}=-\delta_{ij}$ . Here  $\delta$  has the usual meaning of the Kronecker symbol. When integrating over  $k_{0}$  in Eq. (5) we replace the electron mass m by  $m-i\eta$  and take the limit  $\eta \to 0$  from above. This is the same notation and the same way of integration as the one used by Feynman in his original papers in Phys. Rev. 76, 749 (1949); ibid 76, 769 (1949). The natural units ( $\hbar = c = 1$ ) are used throughout the paper. Symbol  $\alpha$ stands for the fine structure constant. In actual calculation we took the value 1/137.03599911 given in [1].
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